

Quantification of model bias underlying the phenomenon of “Einstein from noise”

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Abstract: “Einstein from noise” is an interesting phenomenon arising in cryo-electron microscopy image analysis where spurious patterns could easily emerge by averaging a large number of white-noise images aligned to a reference image through rotation and translation. While this phenomenon can reasonably be explained by model bias, no quantitative studies have been performed to characterize such a bias. We consider a simple framework under which an image is treated as a vector of dimension p and a white-noise image is a random vector uniformly sampled from the $(p-1)$ -dimensional unit sphere. The cross correlation of two images is defined as the inner product of the two corresponding vectors. This framework geometrically explains how the bias results from averaging a properly chosen set of white-noise images that are most highly cross-correlated with the reference image. We quantify the bias in terms of three parameters: the number of white-noise images (n), the image dimension (p), and the size of the selection set (m). Under the conditions that n , p and m are all large and $(\ln n)^2/p$ and m/n are both small, we show that the bias is approximately $\sqrt{\frac{2\gamma}{1+2\gamma}}$ where $\gamma = \frac{m}{p} \ln\left(\frac{n}{m}\right)$.

Key words and phrases: Cryogenic electron microscopy; cross correlation; extreme value distribution; model bias; high dimensional data analysis; white-noise image

1. Introduction

2 Cryogenic electron microscopy (Cryo-EM) is an imaging technique that
uses transmitted electron waves to obtain projection images of a biological
4 sample. In contrast to X-ray crystallography, single particle cryo-EM
does not need crystals and thereby is amenable to structural determination
6 of proteins that are refractory to crystallization, including membrane
proteins (Liao et al., 2013) and yeast spliceosomes that exhibit dynamic
8 patterns (Liao et al., 2013; Yan et al., 2015). This capability enables single
particle cryo-EM to record structures in solution. Because of single particle
10 cryo-EM breakthroughs in high-resolution structure determination of
biomolecules in solution, Nature Methods named cryo-EM as the “Method
12 of the Year” in 2016, and the Nobel Prize in Chemistry in 2017 was awarded
to Jacques Dubochet, Joachim Frank, and Richard Henderson for their pioneering
14 works in developing cryo-EM.

When cryo-EM is applied to imaging biomolecules, the data is recorded
16 on a micrograph, which contains many particle projections in unknown ori-

entations. The signal-to-noise ratio (SNR) of cryo-EM images in general is
18 extremely low (SNR < 0.1) because the biomolecules are photographed with
low exposure to minimize structural degradation caused by radiation. The
20 resulting averages of 2D clustering in cryo-EM processing greatly enhance
SNR of many views and allow the clustering averages to be labeled. Yet,
22 meaningful clustering depends on good image alignment, for which all possible
rotations and translations are exhaustively searched to find the most
24 fitted solution (Frank, 1975; Frank and Al-Ali, 1975; Saxton and Frank,
1976).

26 In practice, there have been cases when cryo-EM processing failed to
converge to a true structure. The pitfall would occur when the particles are
28 small (Mao et al., 2013) or image contrast is low (Murray et al., 2013). In
those cases, the processing was dictated by the reference of a model (Stew-
30 art and Grigorieff, 2004). To elucidate the model bias phenomenon, the
Grigorieff group did an experiment by generating 1000 white-noise images
32 and aligning each of them to an Einstein’s facial image through rotation
and translation. A blurred Einstein’s face emerged from averaging the 1000
34 aligned images. Henderson (2013) further dubbed such unwanted outcome
by “Einstein from noise” and used it to warn the community that an incor-
36 rect 3D density map could be constructed when data are blindly fitted to

a model.

38 In a recent review paper, Lai et al. (2020) discussed the “Einstein from
noise” phenomenon from a statistical perspective. To avoid the technical
40 issue of how rotating an image may destroy the pixel format, they consid-
ered a simple mathematical framework under which an image is treated as a
42 vector of dimension p and a white-noise image is a random vector uniformly
distributed on the $(p - 1)$ -dimensional unit sphere. The cross correlation
44 of two images is defined as the inner product of the two corresponding vec-
tors. Under this framework, we present in Section 2 a simulation study
46 with $n = 2 \times 10^6$ white-noise images with the pixel number $p = 120 \times 120$.
Among the 2×10^6 white-noise images, the largest cross correlation value
48 with Einstein’s facial image (the reference) is merely 0.039, while the cross
correlation increases dramatically to 0.650 after averaging the $m = 800$ im-
50 ages that have the largest cross correlation values with Einstein’s facial im-
age. This illustrates the essence of the “Einstein from noise” phenomenon.
52 The objective of the present paper is to provide a thorough study of the
“Einstein from noise” phenomenon based on the statistical perspective laid
54 out in Lai et al. (2020). A main task is to approximate the distribution
of the cross correlation between the (normalized) average of the m selected
56 images and the reference, which is referred to the (image selection) bias.

While the bias depends on the three parameters n , p , and m in a convoluted
58 manner, under the conditions that n , p and m are all large and $(\ln n)^2/p$
and m/n are both small, we show that the bias is approximately $\sqrt{\frac{2\gamma}{1+2\gamma}}$
60 where $\gamma = \frac{m}{p} \ln \left(\frac{n}{m}\right)$.

The rest of this paper is organized as follows. Section 2 consists of four
62 parts: (i) introducing notation, terminology and the statistical model; (ii)
demonstrating the phenomenon of “Einstein from noise”; (iii) presenting
64 the asymptotic distribution of the largest cross correlation value as n and p
both tend to infinity; (iv) stating asymptotic results on the bias as n , p , and
66 m all tend to infinity. The theoretical results in part (iv) are validated via
simulation as presented in Section 3. Section 4 contains concluding remarks.
68 Proofs of Theorems 1-4 in Section 2 are relegated to the Appendix. The
online supplementary material contains the proofs of auxiliary lemmas.

70 2. Statistical Model and Main Results

2.1 Notation, terminology, and model

72 Let \mathbf{R} be the reference matrix (the digital version of the reference image)
of dimension $d_1 \times d_2$. We assume that $\|\mathbf{R}\| = 1$ where $\|\cdot\|$ denotes the
74 Frobenius norm of a matrix or Euclidean norm of a vector. We generate n
independent and identically distributed (iid) white-noise images as follows.

2.1 Notation, terminology, and model

Let $\mathbf{Z}_1, \dots, \mathbf{Z}_n$ be iid $d_1 \times d_2$ random matrices such that the $d_1 d_2$ components of each \mathbf{Z}_i are iid standard normal. We refer to $\mathbf{Z}_i / \|\mathbf{Z}_i\|$, $i = 1, \dots, n$ (the normalized version of \mathbf{Z}_i) as n iid white-noise images.

Let $\mathbf{r} = \text{vec}(\mathbf{R})$, the p -dimensional column vector which is the vectorized version of \mathbf{R} , where $p = d_1 d_2$. The fact that $\|\mathbf{r}\| = 1$ implies $\mathbf{r} \in \mathcal{S}^{p-1}$ (the $(p-1)$ -dimensional unit sphere). Let $\mathbf{X}_i = \text{vec}(\mathbf{Z}_i) / \|\mathbf{Z}_i\|$. Thus, $\mathbf{X}_1, \dots, \mathbf{X}_n$ are iid uniformly distributed on \mathcal{S}^{p-1} . We refer to both $\mathbf{Z}_i / \|\mathbf{Z}_i\|$ and \mathbf{X}_i as the i -th white-noise image. The cross correlation of \mathbf{X}_i and \mathbf{r} (or equivalently $\mathbf{Z}_i / \|\mathbf{Z}_i\|$ and \mathbf{R}) is defined as $\mathbf{r}^\top \mathbf{X}_i$ (the inner product of \mathbf{X}_i and \mathbf{r}), where \mathbf{r}^\top denotes the transpose of \mathbf{r} . Note that $\mathbf{r}^\top \mathbf{X}_i = \cos \Theta_i$, where Θ_i is the angle between \mathbf{r} and \mathbf{X}_i .

The n white-noise images are ordered (and denoted by $\mathbf{X}^{(1)}, \dots, \mathbf{X}^{(n)}$) according to their cross correlation values with \mathbf{r} . In other words, $(\mathbf{X}^{(1)}, \dots, \mathbf{X}^{(n)})$ is a permutation of $(\mathbf{X}_1, \dots, \mathbf{X}_n)$ such that $\mathbf{r}^\top \mathbf{X}^{(1)} \geq \mathbf{r}^\top \mathbf{X}^{(2)} \geq \dots \geq \mathbf{r}^\top \mathbf{X}^{(n)}$. Let $\Theta_{1:n} \leq \Theta_{2:n} \leq \dots \leq \Theta_{n:n}$ be the order statistics of the angles $(\Theta_1, \dots, \Theta_n)$, so that $\cos \Theta_{i:n} = \mathbf{r}^\top \mathbf{X}^{(i)}$, $i = 1, \dots, n$. Let $\overline{\mathbf{X}}_m = m^{-1} \sum_{i=1}^m \mathbf{X}^{(i)}$. Then $\overline{\mathbf{X}}_m / \|\overline{\mathbf{X}}_m\| \in \mathcal{S}^{p-1}$ is the normalized average of the m white-noise images that are most highly cross-correlated with the reference image. Our goal is to find a good approximation of the distribution of $\rho_{n,p,m} = \mathbf{r}^\top \overline{\mathbf{X}}_m / \|\overline{\mathbf{X}}_m\|$ when n , p , and m are large. Note that for $m = 1$, $\rho_{n,p,1} = \mathbf{r}^\top \mathbf{X}^{(1)} = \cos \Theta_{1:n}$,

2.2 Demonstration of the “Einstein from noise” phenomenon

96 is the largest cross correlation value.

2.2 Demonstration of the “Einstein from noise” phenomenon

98 We now present two figures summarizing the simulation study described
in Section 1, where $n = 2 \times 10^6$, $p = d_1 \times d_2 = 120 \times 120 = 14400$,
100 and $m = 100, 200, 400, 800$. In Figure 1, the leftmost (reference) image
is Einstein’s face, and the other 4 images correspond to $\bar{\mathbf{X}}_m / \|\bar{\mathbf{X}}_m\|$ for
102 $m = 1, 200, 400, 800$. The second image from the left corresponds to $\mathbf{X}^{(1)}$,
whose cross correlation (CC) value with Einstein’s facial image is 0.039
104 (which is the largest among the 2×10^6 white-noise images generated in
the simulation). While this image is rather noisy, Einstein’s face emerges
106 in the other 3 images with different degrees of blurring, corresponding to
CC values 0.426, 0.536, and 0.650.

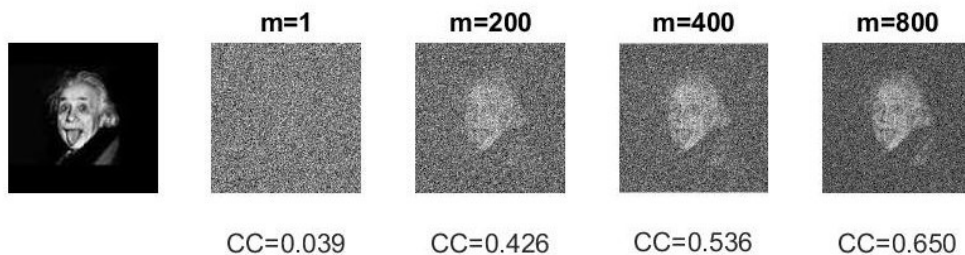


Figure 1: Example with Einstein’s face as the reference image.

2.2 Demonstration of the “Einstein from noise” phenomenon

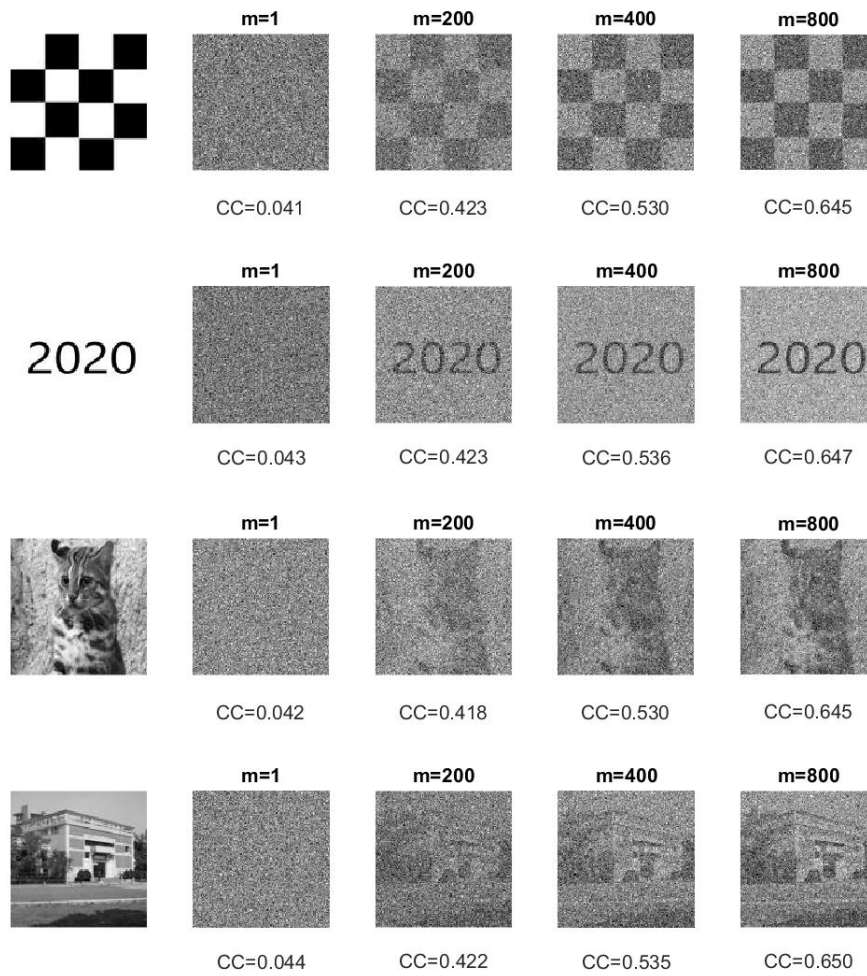


Figure 2: The phenomenon of “Einstein from noise” is shown across various reference images.

2.3 Asymptotic distribution of the largest cross correlation

108 Figure 2 shows similar results with four different reference images of a
 simple chessboard, digits of 2020, a leopard cat and Statistics Building of
 110 Academia Sinica, indicating that the phenomenon of “Einstein from noise”
 is robust across various reference images. The cross correlation values in
 112 Figure 2 are about the same across different reference images, which can be
 explained by the fact that if \mathbf{X} is uniformly distributed on \mathcal{S}^{p-1} , then the
 114 distribution of $\mathbf{r}^\top \mathbf{X}$ is independent of \mathbf{r} .

116 2.3 Asymptotic distribution of the largest cross correlation

Recall that $\cos \Theta_{1:n}$ is the largest cross correlation. The following theorem
 118 provides an approximation to the distribution of $\cos \Theta_{1:n}$ when n and p are
 large.

120 **Theorem 1.** *Let*

$$K_{n,p} = -\ln n + \frac{1}{2} \ln \ln n - \frac{1}{2} \ln \left(\frac{2^{\frac{\ln n}{p}}}{1 - \exp\left(-2^{\frac{\ln n}{p}}\right)} \right) + \frac{1}{2} \ln(4\pi). \quad (1)$$

We have

$$(p-1) \ln(\sin \Theta_{1:n}) - K_{n,p} \xrightarrow{d} G \text{ uniformly as } n \wedge p \rightarrow \infty, \quad (2)$$

122 where $n \wedge p = \min\{n, p\}$, \xrightarrow{d} denotes convergence in distribution, and the

2.4 Asymptotic results on $\rho_{n,p,m}$

cumulative distribution function of G is given by $G(t) = 1 - e^{-e^t}$, $t \in \mathbb{R}$,

124 which is known as the extreme value distribution of Gumbel type.

Based on (2), for $0 < \alpha < 1$, the approximate 100α -th quantile of the distribution of $\cos \Theta_{1:n}$ is

$$M_{n,p}(\alpha) = \sqrt{1 - \exp\{2(K_{n,p} + \ln \ln \alpha^{-1})/(p-1)\}}.$$

Recall that $\cos \Theta_{1:n} = 0.039$ in the simulation study summarized in Figure

126 1, where $n = 2 \times 10^6$ and $p = 120 \times 120$. This observed value is compatible with the approximate 10th quantile $M_{n,p}(0.1) = 0.039$.

128 Figure 3 plots $M_{n,p}(\alpha)$ versus $\log_{10} n$ for $n \leq 10^{100}$ with $p = 120 \times 120$ and $\alpha = .05, .5, .95$. Note that the three quantile curves are very close to each other, indicating that $\cos \Theta_{1:n}$ has a small standard deviation. Figure 130 3 suggests that for $P(\cos \Theta_{1:n} \geq 0.1)$ to be at least 0.05, n is required to be greater than 10^{30} , and for $P(\cos \Theta_{1:n} \geq 0.15)$ to be at least 0.05, n is required to be greater than 10^{70} . In other words, it is unlikely for any of 132 n iid white-noise images of dimension 120×120 to have a cross correlation value with Einstein's face greater than 0.15 unless n is astronomically large.

136 2.4 Asymptotic results on $\rho_{n,p,m}$

When $p = p_n$ and $m = m_n$ both grow with n , asymptotic expansions for the distribution of $\rho_{n,p,m}$ are more involved. Our analysis requires the condition

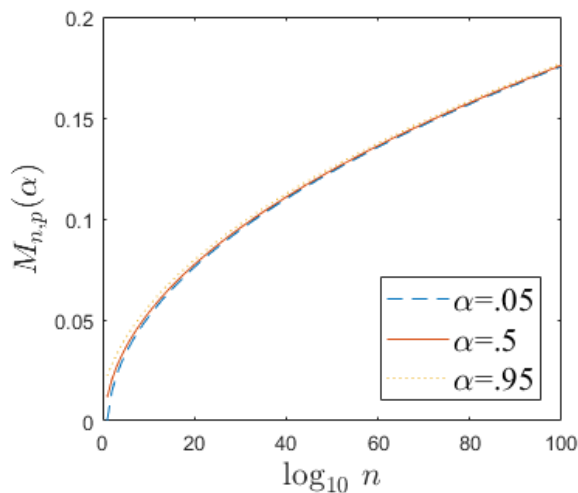


Figure 3: The approximate 100α -th quantile of the distribution of $\cos \Theta_{1:n}$ ($M_{n,p}(\alpha)$) versus $\log_{10} n$ with $p = 120 \times 120$, $\alpha = .05, .5, .95$.

$(\ln n)^2/p = o(1)$ (which is stronger than $(\ln n)/p = o(1)$), so that terms such as $(\ln n)(\ln \ln n)/p$ become negligible. Let

$$\beta_{n,p,m} = \frac{m}{p} \left\{ 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 \right\},$$

which is a model bias index.

138 **Theorem 2.** *Let $p = p_n \rightarrow \infty$ satisfy $(\ln n)^2/p = o(1)$ and $m = m_n \rightarrow \infty$ satisfy $m/n = o(1)$. Then*

$$\rho_{n,p,m}^2 = \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} (1 + o_p(1)).$$

140 *Consequently, $\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \rightarrow 0$ in probability.*

Theorem 3. Let $p = p_n \rightarrow \infty$ satisfy $(\ln n)^2/p = o(1)$ and $m = m_n \rightarrow \infty$

142 satisfy $m(\ln \ln n)^4/(\ln n)^2 = o(1)$. Then

$$\alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \xrightarrow{d} N(0, 1),$$

where $\alpha_{n,p,m} = p (8m + 2p \beta_{n,p,m}^2)^{-1/2} (1 + \beta_{n,p,m})^2$ and $N(0, 1)$ denotes the

144 standard normal distribution.

Theorem 4. Let $p = p_n \rightarrow \infty$ and $m = m_n \rightarrow \infty$.

(i) If $(\ln n)^2/p = o(1)$ and $m/n = o(1)$, then

$$\frac{\rho_{n,p,m}}{\sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}} = 1 + o_p(1).$$

Consequently,

$$\rho_{n,p,m} = \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} + o_p(1) \quad \text{and} \quad E(\rho_{n,p,m}) = \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} + o(1).$$

(ii) In addition to the conditions specified in (i), if $m(\ln \ln n)^4/(\ln n)^2 = o(1)$,

146 then

$$\tilde{\alpha}_{n,p,m} \left(\rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \right) \xrightarrow{d} N(0, 1),$$

where $\tilde{\alpha}_{n,p,m} = 2\alpha_{n,p,m} \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}$.

148 **Remark 1.** On top of the condition $(\ln n)^2/p = o(1)$, Theorem 2 only

requires the mild condition $m/n = o(1)$. Let $\gamma_{n,p,m} = \frac{m}{p} \ln \frac{n}{m}$. Since

150 $\beta_{n,p,m} = 2\gamma_{n,p,m}(1+o(1))$ (i.e. $2\gamma_{n,p,m}$ is the leading term of $\beta_{n,p,m}$), Theorem 2 implies

$$\rho_{n,p,m}^2 = \frac{2\gamma_{n,p,m}}{1 + 2\gamma_{n,p,m}} + o_p(1).$$

152 Consequently,

$$\rho_{n,p,m} = \sqrt{\frac{2\gamma_{n,p,m}}{1+2\gamma_{n,p,m}}} + o_p(1) \quad \text{and} \quad E(\rho_{n,p,m}) = \sqrt{\frac{2\gamma_{n,p,m}}{1+2\gamma_{n,p,m}}} + o(1). \quad (3)$$

Remark 2. To establish asymptotic normality of $\rho_{n,p,m}^2$ (and $\rho_{n,p,m}$), Theorem 3 (and Theorem 4) requires the stringent condition $m(\ln \ln n)^4/(\ln n)^2 = o(1)$. It is unclear whether asymptotic normality still holds when m grows at a rate faster than $(\ln n)^2/(\ln \ln n)^4$. It should also be remarked that under the conditions as in Theorem 3, it is not true that $\alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{2\gamma_{n,p,m}}{1+2\gamma_{n,p,m}} \right) \xrightarrow{d} N(0, 1)$. This shows that while $2\gamma_{n,p,m}$ is the leading term of $\beta_{n,p,m}$, the remaining terms also play a non-negligible role in the proof of asymptotic normality.

Remark 3. Fan et al. (2018) developed an asymptotic theory to approximate the distribution of the maximum spurious correlation of a response variable Y with the best m linear combinations of p covariates \mathbf{X} based on an iid sample of size n when \mathbf{X} and Y are independent. See also Fan et al. (2012) for related results. In our setting, the quantity $\rho_{n,p,m}$ may be referred to as the spurious cross correlation of the reference with the

normalized average of the m white-noise images that are most highly cross-
 168 correlated with the reference. Indeed, with the roles of n and p reversed,
 $\rho_{n,p,m}$ corresponds to another spurious correlation of the response variable
 170 Y with the the average of the m (standardized) covariates in \mathbf{X} that are
 most highly correlated with Y when the p covariates in \mathbf{X} and Y are all
 172 mutually independent.

3. Simulation Results on $\rho_{n,p,m}$

174 By Theorem 4(i), if m is small compared to n and $(\ln n)^2$ is small compared
 to p , then $E(\rho_{n,p,m})$ is expected to be close to $\sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$ while the standard
 176 deviation (s.d.) of $\rho_{n,p,m}$ is expected to be small. We conducted a simulation
 study of the distribution of $\rho_{n,p,m}$ for various combinations of (n, p, m) with
 178 $n = 10^4, 10^5$, $p = 10^4, 4 \times 10^4$, and $m = 100, 200, 400, 600$. The results are
 reported in Tables 1 and 2 where $E(\rho_{n,p,m})$ and $s.d.(\rho_{n,p,m})$ were estimated
 180 based on 1000 replications for each case. While $\sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$ approximates
 $E(\rho_{n,p,m})$ well, it slightly overestimates $E(\rho_{n,p,m})$, more notably for $n = 10^4$.
 182 Clearly, $E(\rho_{n,p,m})$ increases as n or m increases or p decreases. On the
 other hand, $s.d.(\rho_{n,p,m})$ is small ($< .005$) in all cases. Besides, $s.d.(\rho_{n,p,m})$
 184 decreases as n or p increases, and is about the same as m varies from 100 to
 600. Also included in Tables 1 and 2 are $\tilde{\alpha}_{n,p,m}^{-1}$ and the empirical probability

186 (denoted as Prob.) that

$$\left| \rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \right| < 1.96 \tilde{\alpha}_{n,p,m}^{-1}.$$

It is clear from the tables that $\tilde{\alpha}_{n,p,m}^{-1}$ approximates s.d. $(\rho_{n,p,m})$ reasonably well in all cases. By Theorem 4(ii), the Prob. value is expected to be close to .95 if the normal approximation is accurate. By Theorems 3 and 4, $\alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right)$ and $\tilde{\alpha}_{n,p,m} \left(\rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \right)$ are approximately standard normal under somewhat stringent conditions on the growth rates of m and p as $n \rightarrow \infty$. While none of the combinations of (n, p, m) with $n = 10^4, 10^5$, $p = 10^4, 4 \times 10^4$ and $m = 100, 200, 400, 600$ seems to satisfy the condition that $m (\ln \ln n)^4 / (\ln n)^2$ be small, the normal approximation appears to be acceptable for $n = 10^5$ but less satisfactory for $n = 10^4$.

Table 1: $p = 10^4$.

m	$n = 10^4$				$n = 10^5$			
	100	200	400	600	100	200	400	600
$E(\rho_{n,p,m})$	0.257	0.323	0.395	0.437	0.318	0.408	0.509	0.570
$\sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$	0.258	0.325	0.399	0.442	0.319	0.409	0.510	0.571
s.d. $(\rho_{n,p,m})$	0.0043	0.0045	0.0046	0.0048	0.0039	0.0039	0.0040	0.0037
$\tilde{\alpha}_{n,p,m}^{-1}$	0.0051	0.0053	0.0055	0.0057	0.0041	0.0042	0.0040	0.0039
Prob.	0.974	0.967	0.942	0.870	0.967	0.959	0.947	0.953

Table 2: $p = 4 \times 10^4$.

m	$n = 10^4$				$n = 10^5$			
	100	200	400	600	100	200	400	600
$E(\rho_{n,p,m})$	0.132	0.168	0.210	0.236	0.165	0.218	0.283	0.327
$\sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$	0.132	0.169	0.212	0.239	0.166	0.219	0.284	0.328
s.d. $(\rho_{n,p,m})$	0.0022	0.0024	0.0026	0.0027	0.0019	0.0020	0.0021	0.0022
$\tilde{\alpha}_{n,p,m}^{-1}$	0.0026	0.0028	0.0031	0.0033	0.0021	0.0022	0.0023	0.0023
Prob.	0.977	0.978	0.946	0.871	0.968	0.967	0.955	0.953

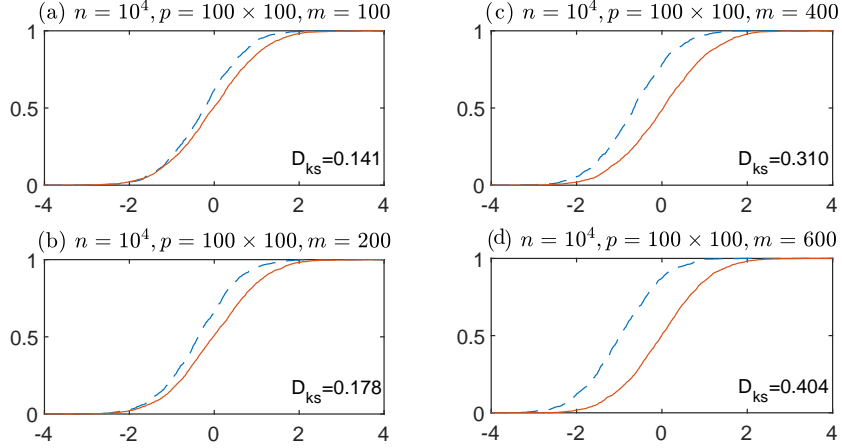


Figure 4: Empirical pdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf. (solid curves): $n = 10^4$, $p = 10^4$.

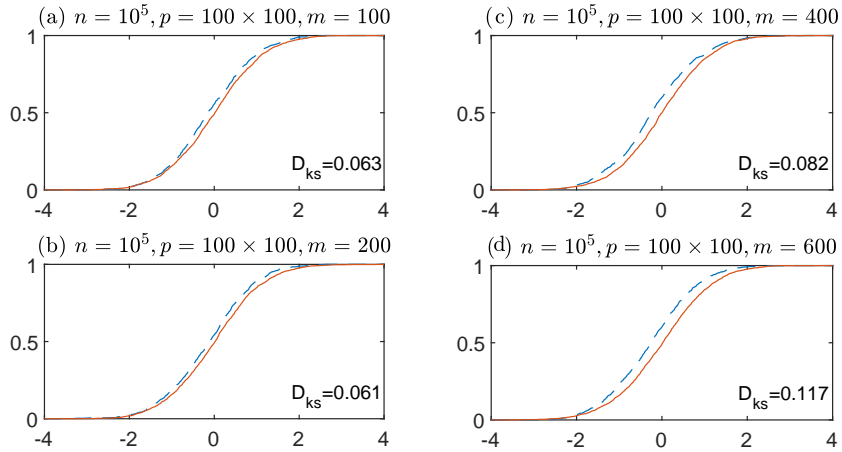


Figure 5: Empirical pdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf. (solid curves): $n = 10^5$, $p = 10^4$.

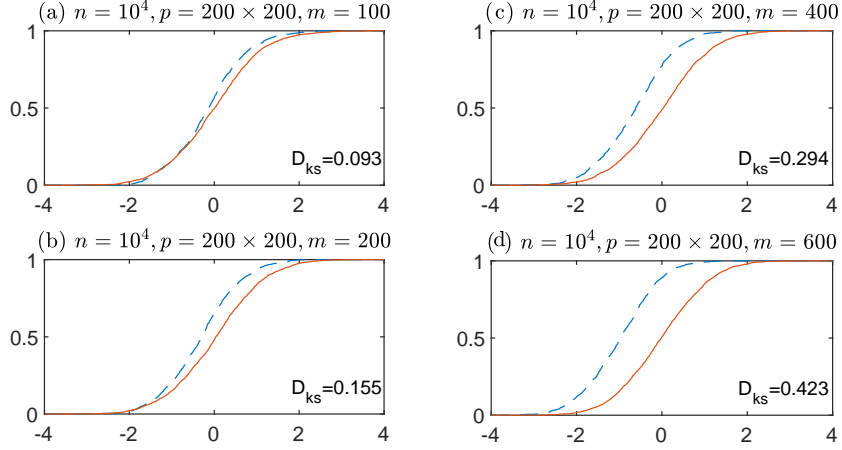


Figure 6: Empirical pdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf. (solid curves): $n = 10^4$, $p = 4 \times 10^4$.

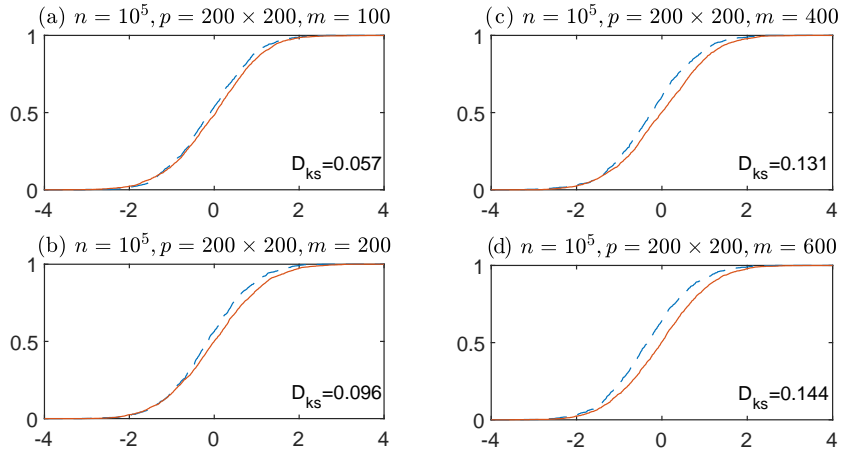


Figure 7: Empirical pdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf. (solid curves): $n = 10^5$, $p = 4 \times 10^4$.

In Figures 4-7, we plot the empirical cumulative distribution function
 (cdf) of $\tilde{\alpha}_{n,p,m} \left(\rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}} \right)$ (based on 1000 replications), along with
 the standard normal cdf for each combination of (n, p, m) . (The value of D_{ks}
 is the Kolmogorov-Smirnov distance between the two cdfs.) The empirical
 cdf is shifted to the left of the standard normal cdf (more notably for $n =$
 10^4 in Figures 4 and 6), indicating that the mean of $\rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$
 is negative. This is consistent with the results in Tables 1 and 2 where
 $\sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$ (slightly) overestimates $E(\rho_{n,p,m})$ (more notably for $n = 10^4$).

4. Concluding Remarks

This paper studied a simple statistical model in order to quantitatively ex-
 amine the phenomenon of “Einstein from noise”. Specifically, for a given
 reference image of dimension p and a set S_n of n iid white-noise images
 (with the common uniform distribution on \mathcal{S}^{p-1}), we derived the asymp-
 totic behavior of the cross correlation $\rho_{n,p,m}$ between the reference and the
 normalized average of the m “most biased” members in S_n in the sense
 that they have the largest cross correlation values with the reference. Our
 theoretical results indicate that for $m = 1$ and $p = 120 \times 120$, unless n is far
 beyond the practical range ($> 10^{70}$), $\rho_{n,p,1}$ is small (< 0.15) with high prob-
 ability, implying that none of n white-noise images even remotely resembles

216 the reference. On the other hand, for m moderately large (≥ 400), $\rho_{n,p,m}$
exceeds 0.5 with high probability if $n = 2 \times 10^6$, in which case a blurred
218 version of the reference emerges from the normalized average of the m most
biased members in S_n .

220 Given a set S_n of n iid white-noise images, Cai et al. (2013) derived the
asymptotic distribution of the maximum of all pairwise cross correlations
222 in S_n . See also Cai and Jiang (2011, 2012) and references therein. In
the absence of a reference image, their results may be applied to test the
224 null hypothesis that S_n consists of n iid white-noise images. On the other
hand, given a reference image, our results can be used to test such a null
226 hypothesis against the alternative that some of the n images in S_n are biased
towards the reference by checking whether $\rho_{n,p,m}$ exceeds a threshold (which
228 is determined by the null distribution of $\rho_{n,p,m}$).

Our approach can be directly generalized to tackle a special case of
230 multiple references. Let $\mathbf{r}^{(1)}, \dots, \mathbf{r}^{(k)}$ be k given references of dimension
 p . Given a set S_n of n iid white-noise images, for $i = 1, \dots, k$, let $\rho_{n,p,m}^{(i)}$
232 ($i = 1, \dots, k$) denote the cross correlation between $\mathbf{r}^{(i)}$ and the normal-
ized average of those m members in S_n having the largest cross correlation
234 values with $\mathbf{r}^{(i)}$. It would be of interest to derive the asymptotic distri-
bution of $\max\{\rho_{n,p,m}^{(i)} : i = 1, \dots, k\}$. If $\mathbf{r}^{(1)}, \dots, \mathbf{r}^{(k)}$ are orthogonal (i.e.

236 the pairwise cross correlations are all equal to 0), then it can be argued
that $\rho_{n,p,m}^{(1)}, \dots, \rho_{n,p,m}^{(k)}$ are asymptotically independent, so that the asymp-
238 totic distribution of $\max\{\rho_{n,p,m}^{(i)} : i = 1, \dots, k\}$ can be readily derived by
Theorem 4. However, it seems difficult to find the asymptotic distribution
240 of $\max\{\rho_{n,p,m}^{(i)} : i = 1, \dots, k\}$ when $\mathbf{r}^{(1)}, \dots, \mathbf{r}^{(k)}$ are not orthogonal.

The phenomenon of “Einstein from noise” originally arose in the con-
242 text of cryo-EM image analysis where a key component is image alignment
(including rotation and translation). While to address this more compli-
244 cated problem is beyond the scope of the present paper, it is worth noting
that the geometric shape of the reference is likely to play a significant role
246 in the asymptotic theory yet to be developed. As an example, consider a
rotationally invariant reference, e.g. an image of a centered wheel. Because
248 of rotational symmetry of the reference, a data image cannot fit the refer-
ence any better by rotation. We leave this challenging problem for future
250 work.

Supplementary Material

252 The online Supplementary Material contains the proofs of Lemmas A6-
A8 stated in the Appendix.

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258 A. Appendix

The Appendix consists of three sections. Section 4 states some auxiliary
 260 lemmas, Section 4 contains the proof of Theorem 1, and Section 4 provides
 the proofs of Theorems 2-4. For easy reference, a complete list of notations
 262 is given in Supplementary Material. Note that if \mathbf{X} is uniformly distributed
 on \mathcal{S}^{p-1} , then the distribution of $\mathbf{r}^\top \mathbf{X}$ is the same for all $\mathbf{r} \in \mathcal{S}^{p-1}$. Without
 264 loss of generality, we assume $\mathbf{r} = (1, 0, \dots, 0)^\top \in \mathcal{S}^{p-1}$.

A.1. Auxiliary lemmas

266 **Lemma A1.** (*Lemma 6.2 of Cai and Jiang (2012)*) For $t \in (0, 1)$, we have

$$\left(1 + \frac{1}{pt^2}\right)^{-1} \frac{1}{(p+2)t} (1-t^2)^{(p+2)/2} \leq \int_t^1 (1-u^2)^{p/2} du \leq \frac{1}{(p+2)t} (1-t^2)^{(p+2)/2}.$$

Since \mathbf{X}_i , $i = 1, \dots, n$ are iid uniformly distributed on \mathcal{S}^{p-1} and Θ_i
 268 denotes the angle between \mathbf{X}_i and $\mathbf{r} = (1, 0, \dots, 0)^\top$, we have (cf. Eq (5))

of Cai et al. (2013)) that $\Theta_i, i = 1, \dots, n$ are iid with the common cdf

$$\begin{aligned} F_p(\theta) &= \int_0^\theta \frac{1}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} (\sin x)^{p-2} dx \\ &= \int_{\cos \theta}^1 \frac{1}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} (1-u^2)^{\frac{p-3}{2}} du, \quad \theta \in [0, \pi]. \end{aligned} \quad (\text{A.1})$$

270 Let

$$\bar{F}_p(\theta) = \frac{1}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} \frac{\sin^{p-1} \theta}{(p-1)|\cos \theta|}. \quad (\text{A.2})$$

The following lemma is a consequence of Lemma A1.

Lemma A2. *For $\theta \in (0, \pi/2)$ and $p > 3$, we have*

$$\left(1 + \frac{1}{(p-3)\cos^2 \theta}\right)^{-1} \bar{F}_p(\theta) \leq F_p(\theta) \leq \bar{F}_p(\theta).$$

272 Let U_1, U_2, \dots be iid uniform (0,1) random variables and let $U_{1:n} \leq \dots \leq U_{n:n}$ denote the order statistics of U_1, \dots, U_n . Let $S_0 = 0$, and
 274 $S_i = \xi_1 + \dots + \xi_i, i = 1, 2, \dots$, where ξ_1, ξ_2, \dots are iid exponential random variables with mean 1. The next lemma is well known; see e.g. Karlin and
 276 Taylor (1975). We write $\mathbf{X} \stackrel{d}{=} \mathbf{Y}$ if random vectors \mathbf{X} and \mathbf{Y} are equal in distribution.

278 **Lemma A3.** (i) $(U_{1:n}, \dots, U_{n:n}) \stackrel{d}{=} (S_1, \dots, S_n)/S_{n+1}$. (ii) $(S_1, \dots, S_n)/S_{n+1}$ is independent of S_{n+1} .

280 Recall that $(\mathbf{X}^{(1)}, \dots, \mathbf{X}^{(n)})$ is a permutation of $(\mathbf{X}_1, \dots, \mathbf{X}_n)$ such that $X_1^{(1)} \leq \dots \leq X_1^{(n)}$, where $X_1^{(i)} = \mathbf{r}^\top \mathbf{X}^{(i)}$ (the first component of

282 $\mathbf{X}^{(i)}$. Let \mathbf{V}_i and $\mathbf{V}^{(i)}$ be defined by $\mathbf{X}_i = (X_{i1}, (1 - X_{i1}^2)^{1/2} \mathbf{V}_i^\top)^\top$ and
 $\mathbf{X}^{(i)} = (X_1^{(i)}, \nu_i \mathbf{V}^{(i)\top})^\top$, where $\nu_i = (1 - X_1^{(i)2})^{1/2}$. In other words, \mathbf{V}_i ($\mathbf{V}^{(i)}$,
284 respectively) $\in \mathcal{S}^{p-2}$ is the normalized subvector of \mathbf{X}_i ($\mathbf{X}^{(i)}$, respectively)
with the first component deleted.

286 **Lemma A4.**

(i) X_{i1} and $\mathbf{V}_i, i = 1, \dots, n$ are all independent.

288 (ii) $X_{i1}, i = 1, \dots, n$ are iid.

(iii) $\mathbf{V}_i, i = 1, \dots, n$ are iid with the uniform distribution on \mathcal{S}^{p-2} .

290 (iv) $(\mathbf{V}^{(1)}, \dots, \mathbf{V}^{(n)})$ is independent of (X_{11}, \dots, X_{n1}) and hence independent of $(X_1^{(1)}, \dots, X_1^{(n)})$.

292 (v) $\mathbf{V}^{(i)}, i = 1, \dots, n$ are iid with the uniform distribution on \mathcal{S}^{p-2} .

Recall that

$$\bar{\mathbf{X}}_m = \frac{1}{m} \sum_{i=1}^m \mathbf{X}^{(i)} = \left(m^{-1} \sum_{i=1}^m X_1^{(i)}, m^{-1} \sum_{i=1}^m \nu_i \mathbf{V}^{(i)\top} \right)^\top$$

294 and that

$$\rho_{n,p,m}^2 = \left(\mathbf{r}^\top \frac{\bar{\mathbf{X}}_m}{\|\bar{\mathbf{X}}_m\|} \right)^2 = \frac{\left(\frac{1}{m} \sum_{i=1}^m X_1^{(i)} \right)^2}{\left(\frac{1}{m} \sum_{i=1}^m X_1^{(i)} \right)^2 + \left\| \frac{1}{m} \sum_{i=1}^m \nu_i \mathbf{V}^{(i)} \right\|^2}.$$

Let $\mathbf{V}'_i, i = 1, \dots, n$ be iid uniformly distributed on \mathcal{S}^{p-2} and independent
296 of $\mathbf{X}_1, \dots, \mathbf{X}_n$. Then the following lemma is a consequence of Lemma A4.

Lemma A5.

$$\begin{aligned} \rho_{n,p,m}^2 &\stackrel{d}{=} \frac{\left(m^{-1} \sum_{i=1}^m X_1^{(i)}\right)^2}{\left(m^{-1} \sum_{i=1}^m X_1^{(i)}\right)^2 + \left\|m^{-1} \sum_{i=1}^m \nu_i \mathbf{V}'_i\right\|^2} \\ &= \frac{A_{n,p,m}}{A_{n,p,m} + V_{n,p,m}}, \end{aligned} \quad (\text{A.3})$$

where

$$A_{n,p,m} = \left(\frac{1}{m} \sum_{i=1}^m X_1^{(i)}\right)^2 \quad \text{and} \quad V_{n,p,m} = \left\|\frac{1}{m} \sum_{i=1}^m \nu_i \mathbf{V}'_i\right\|^2. \quad (\text{A.4})$$

298 The long proofs of Lemmas A6-A8 below are given in Supplementary Material.

300 **Lemma A6.** *Let $m = m_n \rightarrow \infty$ satisfy $m/n = o(1)$ and $p = p_n \rightarrow \infty$ satisfy $(\ln n)^2/p = O(1)$. Then*

(i)

$$\max_{1 \leq i \leq m} \left| p \ln(\sin \Theta_{i:n}) + \ln \frac{n}{i} - \frac{1}{2} \ln \ln \frac{n}{i} \right| = O_p(1),$$

(ii)

$$\max_{1 \leq i \leq m} \left| -\frac{p}{2} \cos^2 \Theta_{i:n} + \ln \frac{n}{i} - \frac{1}{2} \ln \ln \frac{n}{i} \right| = O_p(1),$$

302 where $\Theta_{1:n} \leq \Theta_{2:n} \leq \dots \leq \Theta_{n:n}$ are the order statistics of $\Theta_1, \dots, \Theta_n$.

Lemma A7. *Suppose that $p = p_n \rightarrow \infty$ satisfies $(\ln n)^2/p = O(1)$.*

(i) *If $m = m_n \rightarrow \infty$ satisfies $m/n \rightarrow 0$, then*

$$-pA_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} = O_p(1).$$

304 (ii) If $m = m_n \rightarrow \infty$ satisfies $(\ln m)^3/(\ln n)^2 \rightarrow 0$, then

$$-pA_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 - \frac{2}{p} \left(\ln \frac{n}{m} \right)^2 = o_p(1).$$

(iii) If $m = m_n \rightarrow \infty$ satisfies $m(\ln \ln n)^4/(\ln n)^2 \rightarrow 0$, then

$$\left(\frac{m}{8} \right)^{1/2} \left\{ -pA_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 - \frac{2}{p} \left(\ln \frac{n}{m} \right)^2 \right\} \xrightarrow{d} N(0, 1).$$

306 **Lemma A8.** Let $\mathbf{W}_1, \dots, \mathbf{W}_n$ be iid uniformly distributed on \mathcal{S}^{p-1} . Then

$$\sqrt{\frac{p}{2n^2}} \sum_{1 \leq i \neq \ell \leq n} \langle \mathbf{W}_i, \mathbf{W}_\ell \rangle \xrightarrow{d} N(0, 1) \text{ uniformly as } n \wedge p \rightarrow \infty,$$

where $\langle \mathbf{W}_i, \mathbf{W}_\ell \rangle$ denotes the inner product of \mathbf{W}_i and \mathbf{W}_ℓ .

308 A.2. Proof of Theorem 1

Theorem 1 is a special case of Theorem A1 below for $m = 1$.

310 **Theorem A1.** Let

$$T_{n,p} = (p-1) \ln(\sin \Theta_{m:n}) - K_{n,p},$$

where $K_{n,p}$ is defined as in (1). Let $G_m^*(t) = G_m(e^t)$, $t \in \mathbb{R}$, where G_m

312 denotes the gamma distribution with shape parameter m and scale parameter

1. Then for fixed $m = 1, 2, \dots$, $T_{n,p} \xrightarrow{d} G_m^*$ uniformly as $n \wedge p \rightarrow \infty$.

Proof. We claim that

$$T_{n,p} = T_{n,p_n} \xrightarrow{d} G_m^* \tag{A.5}$$

if $p = p_n \rightarrow \infty$ satisfies $\lim_{n \rightarrow \infty} \ln n / p = \alpha \in [0, \infty]$. Assume for now that the claim (A.5) holds. To show that $T_{n,p} \xrightarrow{d} G_m^*$ uniformly as $n \wedge p \rightarrow \infty$, suppose to the contrary that $\limsup_{n \wedge p \rightarrow \infty} \sup_{t \in \mathbb{R}} |\mathbb{P}(T_{n,p} \leq t) - G_m^*(t)| > 0$. Then there exist an $\varepsilon > 0$ and a sequence $\{(n_\ell, p_\ell) : \ell = 1, 2, \dots\}$ such that $\lim_{\ell \rightarrow \infty} n_\ell \wedge p_\ell = \infty$ and

$$\sup_{t \in \mathbb{R}} |\mathbb{P}(T_{n_\ell, p_\ell} \leq t) - G_m^*(t)| > \varepsilon \text{ for } \ell = 1, 2, \dots \quad (\text{A.6})$$

314 Choose an arbitrary subsequence $\{(n_{\ell_k}, p_{\ell_k}) : k = 1, 2, \dots\}$ such that $\lim_{k \rightarrow \infty} \ln n_{\ell_k} / p_{\ell_k} = \alpha \in [0, \infty]$. Then (A.6) contradicts (A.5), implying that

316 $T_{n,p} \xrightarrow{d} G_m^*$ uniformly as $n \wedge p \rightarrow \infty$.

We now prove (A.5). Suppose $p = p_n \rightarrow \infty$ satisfies $\lim_{n \rightarrow \infty} \ln n / p = \alpha \in [0, \infty]$. For fixed m , since $F_p(\Theta_{m:n}) \stackrel{d}{=} U_{m:n}$, we have by Lemma A3

$$\begin{aligned} \mathbb{P}(nF_p(\Theta_{m:n}) \leq e^t) &= \mathbb{P}(nU_{m:n} \leq e^t) = \mathbb{P}\left(n \frac{S_m}{S_{n+1}} \leq e^t\right) \\ &\longrightarrow \mathbb{P}(S_m \leq e^t) = G_m(e^t) = G_m^*(t). \end{aligned} \quad (\text{A.7})$$

For fixed $t > 0$, let $t_n \in [0, 1)$ be such that

$$\frac{p-1}{2} \ln(1 - t_n^2) = \min\{K_{n,p} + t, 0\}.$$

Noting that

$$K_{n,p} = K_{n,p_n} = -(\ln n)(1 + o(1)) \text{ as } n \rightarrow \infty, \quad (\text{A.8})$$

318 we have for large n

$$\frac{p-1}{2} \ln(1-t_n^2) = K_{n,p} + t < 0. \quad (\text{A.9})$$

By Lemma A2,

$$\left(1 + \frac{1}{(p-3)t_n^2}\right)^{-1} \bar{F}_p(\cos^{-1} t_n) \leq F_p(\cos^{-1} t_n) \leq \bar{F}_p(\cos^{-1} t_n),$$

implying that

$$\begin{aligned} \mathbb{P}(nF_p(\Theta_{m:n}) \leq n\bar{F}_p(\cos^{-1} t_n)) &\geq \mathbb{P}(nF_p(\Theta_{m:n}) \leq nF_p(\cos^{-1} t_n)) \\ &\geq \mathbb{P}\left(nF_p(\Theta_{m:n}) \leq \left(1 + \frac{1}{(p-3)t_n^2}\right)^{-1} n\bar{F}_p(\cos^{-1} t_n)\right). \end{aligned} \quad (\text{A.10})$$

320 Recalling $\alpha = \lim_{n \rightarrow \infty} (\ln n)/p$, we claim that for every $\alpha \in [0, \infty]$, as $n \rightarrow \infty$

$$n\bar{F}_p(\cos^{-1} t_n) = e^t + o(1), \quad (\text{A.11})$$

$$pt_n^2 \rightarrow \infty, \quad (\text{A.12})$$

$$\mathbb{P}(\cos \Theta_{m:n} \leq -t_n) \rightarrow 0. \quad (\text{A.13})$$

By (A.7), (A.10), (A.11) and (A.12),

$$\mathbb{P}(\cos \Theta_{m:n} \geq t_n) = \mathbb{P}(nF_p(\Theta_{m:n}) \leq nF_p(\cos^{-1} t_n)) \rightarrow G_m^*(t). \quad (\text{A.14})$$

Furthermore,

$$\begin{aligned}
P(T_{n,p} \leq t) &= P\left(\frac{p-1}{2} \ln(1 - \cos^2 \Theta_{m:n}) - K_{n,p} \leq t\right) \\
&= P(\cos^2 \Theta_{m:n} \geq t_n^2) \quad (\text{by (A.9)}) \\
&= P(\cos \Theta_{m:n} \geq t_n) + P(\cos \Theta_{m:n} \leq -t_n) \\
&\rightarrow G_m^*(t) \quad (\text{by (A.13) and (A.14)}).
\end{aligned}$$

It remains to establish (A.11)-(A.13). Note that by Sterling's formula (see e.g. Tricomi and Erdélyi (1951)),

$$\frac{\Gamma(p/2)}{\Gamma((p-1)/2)} = \sqrt{\frac{p}{2}} \left(1 + O\left(\frac{1}{p}\right)\right) \quad \text{as } p \rightarrow \infty. \quad (\text{A.15})$$

We have

$$\begin{aligned}
\ln(n\bar{F}_p(\cos^{-1} t_n)) &= \ln \left\{ \frac{n}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} \left(\frac{(1-t_n^2)^{p-1}}{(p-1)^2 t_n^2}\right)^{1/2} \right\} \quad (\text{by (A.2)}) \\
&= \ln \left\{ n \left(\frac{(1-t_n^2)^{p-1}}{2\pi p t_n^2}\right)^{1/2} \right\} + O\left(\frac{1}{p}\right) \quad (\text{by (A.15)}) \\
&= \frac{p-1}{2} \ln(1-t_n^2) + \ln n - \frac{1}{2} \ln(pt_n^2) - \frac{1}{2} \ln(2\pi) + O\left(\frac{1}{p}\right) \\
&= K_{n,p} + t + \ln n - \frac{1}{2} \ln(pt_n^2) - \frac{1}{2} \ln(2\pi) + O\left(\frac{1}{p}\right) \quad (\text{by (A.9)}).
\end{aligned} \tag{A.16}$$

By (A.8) and (A.9),

$$\ln(1-t_n^2) = -\frac{2\ln n}{p}(1+o(1)), \tag{A.17}$$

implying that

$$t_n \rightarrow (1 - e^{-2\alpha})^{1/2}, \quad (\text{A.18})$$

322 where $\lim_{n \rightarrow \infty} \ln n/p = \alpha \in [0, \infty]$ and $e^{-\infty} := 0$.

If $\alpha = 0$, we have $t_n \rightarrow 0^+$, so that by (A.17)

$$t_n^2 = \frac{2 \ln n}{p}(1 + o(1)), \quad (\text{A.19})$$

from which it follows that $\ln(pt_n^2) = \ln(2 \ln n) + o(1)$. By the definition of $K_{n,p}$, we have $K_{n,p} = -\ln n + (\ln \ln n)/2 + \ln(4\pi)/2 + o(1)$, so that $K_{n,p} + \ln n - \ln(pt_n^2)/2 - \ln(2\pi)/2 = o(1)$, which together with (A.16) establishes (A.11) for $\alpha = 0$. If $0 < \alpha < \infty$, we have $t_n^2 = 1 - e^{-2\alpha} + o(1)$ (by (A.18)) and $\ln(pt_n^2) = \ln \ln n - \ln \alpha + \ln(1 - e^{-2\alpha}) + o(1)$, so that $K_{n,p} + \ln n - \ln(pt_n^2)/2 - \ln(2\pi)/2 = o(1)$, which together with (A.16) establishes (A.11) for $0 < \alpha < \infty$. If $\alpha = \infty$, we have $t_n \rightarrow 1^-$, so that by the definition of $K_{n,p}$,

$$\begin{aligned} & K_{n,p} + \ln n - \frac{1}{2} \ln(pt_n^2) - \frac{1}{2} \ln(2\pi) \\ &= -\ln n + \frac{1}{2} \ln \ln n - \frac{1}{2} \ln \left(\frac{2 \ln n}{p} \right) + \frac{1}{2} \ln(4\pi) + \ln n - \frac{1}{2} \ln p - \frac{1}{2} \ln(2\pi) + o(1) \\ &= o(1), \end{aligned}$$

which together with (A.16) establishes (A.11) for $\alpha = \infty$.

Next, (A.19) holds for $\alpha = 0$, which implies (A.12). For $0 < \alpha \leq \infty$, it
 326 follows from (A.18) that $t_n \rightarrow (1 - e^{-2\alpha})^{1/2} > 0$, which implies (A.12).

Finally, to prove (A.13), note that

$$\mathbb{P}(\cos \Theta_{m:n} \leq -t_n) \leq \mathbb{P}(\Theta_{m:n} \geq \pi/2) = \mathbb{P}(B(n, 1/2) < m) \rightarrow 0,$$

328 where $B(n, 1/2)$ denotes a binomial random variable with parameters n and
 1/2 (success probability). This establishes (A.13) and completes the proof
 330 of Theorem A1. □

A.3. Proofs of Theorems 2-4

332 We first show that if $m = m_n \rightarrow \infty$ satisfies $m/n \rightarrow 0$ and $p = p_n \rightarrow \infty$
 satisfies $(\ln n)^2/p \rightarrow 0$, then

$$m\sqrt{\frac{p}{2}} \left(V_{n,p,m} - \frac{1}{m} \right) \xrightarrow{d} N(0, 1), \tag{A.20}$$

334 where $V_{n,p,m} = \|\frac{1}{m} \sum_{i=1}^m \nu_i \mathbf{V}'_i\|^2$ with $\nu_i^2 = 1 - \cos^2 \Theta_{i:n}$, and $\mathbf{V}'_1, \dots, \mathbf{V}'_m$
 are iid uniformly distributed on \mathcal{S}^{p-2} , and $(\mathbf{V}'_1, \dots, \mathbf{V}'_m)$ is independent of
 336 (ν_1, \dots, ν_m) .

We have

$$\begin{aligned}
V_{n,p,m} &= \frac{1}{m^2} \sum_{i=1}^m \nu_i^2 \|\mathbf{V}'_i\|^2 + \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} \nu_i \nu_\ell \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle \\
&= \frac{1}{m} + \frac{1}{m^2} \sum_{i=1}^m (\nu_i^2 - 1) + \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} \{1 + (\nu_i \nu_\ell - 1)\} \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle \\
&= \frac{1}{m} + V'_{1,n} + V'_{2,n} + V'_{3,n}, \tag{A.21}
\end{aligned}$$

where

$$\begin{aligned}
V'_{1,n} &= \frac{1}{m^2} \sum_{i=1}^m (\nu_i^2 - 1) = -\frac{1}{m^2} \sum_{i=1}^m \cos^2 \Theta_{i:n}, \\
V'_{2,n} &= \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle, \\
V'_{3,n} &= \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} (\nu_i \nu_\ell - 1) \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle.
\end{aligned}$$

By Lemma A8, we have

$$m \sqrt{\frac{p}{2}} V'_{2,n} \xrightarrow{d} N(0, 1). \tag{A.22}$$

338 It remains to prove

$$mp^{1/2} V'_{i,n} = o_p(1), \quad i = 1, 3. \tag{A.23}$$

By Lemma A6(ii),

$$\max_{1 \leq i \leq m} \cos^2 \Theta_{i:n} = O_p \left(\frac{\ln n}{p} \right),$$

340 implying that $mp^{1/2} V'_{1,n} = O_p \left(\frac{\ln n}{p^{1/2}} \right) = o_p(1)$. To show $mp^{1/2} V'_{3,n} = o_p(1)$,

note that (ν_1, \dots, ν_m) is independent of $(\mathbf{V}'_1, \dots, \mathbf{V}'_m)$ and $E[\langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle \langle \mathbf{V}'_{i'}, \mathbf{V}'_{\ell'} \rangle] =$

342 0 if $i \neq \ell$, $i' \neq \ell'$ and $\{i, \ell\} \neq \{i', \ell'\}$. Also, for $i \neq \ell$, $E\langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle^2 = \int_0^\pi \cos^2(\theta) dF_{p-1}(\theta) = \frac{1}{p-1}$, where F_p is defined as in (A.1). We have

$$\begin{aligned}
EV'_{3,n}{}^2 &= \frac{2}{m^4} \sum_{1 \leq i \neq \ell \leq m} E[(\nu_i \nu_\ell - 1)^2] E\langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle^2 \\
&= \frac{2}{m^4} \sum_{1 \leq i \neq \ell \leq m} E[(\nu_i \nu_\ell - 1)^2] \frac{1}{p-1} \\
&= o\left(\frac{1}{m^2 p}\right), \tag{A.24}
\end{aligned}$$

344 since $|\nu_i| \leq 1$ and $\nu_i \nu_\ell - 1 \rightarrow 0$ in probability uniformly in $1 \leq i \neq \ell \leq m$.

It follows from (A.24) that $mp^{1/2}V'_{3,n} = o_p(1)$. This proves (A.23) and

346 completes the proof of (A.20).

Proof of Theorem 2. Since by (A.3) $\rho_{n,p,m}^2 \stackrel{d}{=} \frac{A_{n,p,m}}{A_{n,p,m} + V_{n,p,m}}$, we have

$$\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \stackrel{d}{=} \frac{A_{n,p,m} - \beta_{n,p,m}/m}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})} + \frac{(1/m - V_{n,p,m})\beta_{n,p,m}}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})}. \tag{A.25}$$

Since $\beta_{n,p,m} = \frac{m}{p} \{2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2\}$, it follows from Lemma

348 A7(i) and (A.20) that

$$p(A_{n,p,m} - \frac{1}{m}\beta_{n,p,m}) = O_p(1), \quad mV_{n,p,m} = 1 + o_p(1), \quad p\beta_{n,p,m}V_{n,p,m} = (2 + o_p(1)) \ln \left(\frac{n}{m}\right).$$

Thus,

$$\begin{aligned}
\frac{A_{n,p,m} - \beta_{n,p,m}/m}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})} &= \frac{p(A_{n,p,m} - \beta_{n,p,m}/m)}{(p\beta_{n,p,m}A_{n,p,m} + p\beta_{n,p,m}V_{n,p,m})(1 + \beta_{n,p,m})} = o_p(1) \frac{\beta_{n,p,m}}{(1 + \beta_{n,p,m})}, \\
\frac{(1/m - V_{n,p,m})\beta_{n,p,m}}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})} &= \frac{(1 - mV_{n,p,m})}{(mA_{n,p,m} + mV_{n,p,m})} \frac{\beta_{n,p,m}}{(1 + \beta_{n,p,m})} = o_p(1) \frac{\beta_{n,p,m}}{(1 + \beta_{n,p,m})}.
\end{aligned}$$

350

We have by (A.25),

$$\rho_{n,p,m}^2 = \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}(1 + o_p(1)).$$

The proof is complete. \square

352

Proof of Theorem 3. By (A.21)-(A.23),

$$\begin{aligned} m\sqrt{\frac{p}{2}} \left(V_{n,p,m} - \frac{1}{m} \right) &= m\sqrt{\frac{p}{2}} (V'_{1,n} + V'_{2,n} + V'_{3,n}) \\ &= m\sqrt{\frac{p}{2}} V'_{2,n} + o_p(1). \end{aligned} \quad (\text{A.26})$$

354 Let

$$\begin{aligned} Z_{1,n} &= p\sqrt{\frac{m}{8}} \left(A_{n,p,m} - \frac{1}{m}\beta_{n,p,m} + \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \right), \\ Z_{2,n} &= m\sqrt{\frac{p}{2}} V'_{2,n}, \\ \gamma_n &= (A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m}). \end{aligned}$$

We have by (A.25) and (A.26)

$$\begin{aligned} &\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \\ \stackrel{d}{=} &\gamma_n^{-1} \left\{ \frac{1}{p\sqrt{m/8}} Z_{1,n} - \frac{\beta_{n,p,m}}{m\sqrt{p/2}} m\sqrt{\frac{p}{2}} \left(V_{n,p,m} - \frac{1}{m} \right) \right\} - \gamma_n^{-1} \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \\ = &\gamma_n^{-1} \left\{ \sqrt{\frac{8}{mp^2}} Z_{1,n} - \sqrt{\frac{2}{m^2 p}} \beta_{n,p,m} (Z_{2,n} + o_p(1)) \right\} - \gamma_n^{-1} \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \\ = &\gamma_n^{-1} \left(\frac{8}{mp^2} + \frac{2}{m^2 p} \beta_{n,p,m}^2 \right)^{1/2} \{c_{1,n} Z_{1,n} + c_{2,n} (Z_{2,n} + o_p(1))\} - \gamma_n^{-1} \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2, \end{aligned} \quad (\text{A.27})$$

356 where

$$\begin{aligned} c_{1,n} &= \sqrt{\frac{8}{mp^2}} \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{-1/2}, \\ c_{2,n} &= -\sqrt{\frac{2}{m^2p}} \beta_{n,p,m} \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{-1/2}. \end{aligned}$$

Since $\rho_{n,p,m}^2 \stackrel{d}{=} A_{n,p,m}/(A_{n,p,m} + V_{n,p,m})$, we have by Theorem 2

$$\frac{A_{n,p,m}}{A_{n,p,m} + V_{n,p,m}} = \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} (1 + o_p(1)). \quad (\text{A.28})$$

358 It follows from Lemma A7(i) and $(p/m)\beta_{n,p,m} = 2 \ln \frac{n}{m} (1 + o(1))$ that

$$\frac{mA_{n,p,m}}{\beta_{n,p,m}} = \frac{pA_{n,p,m}}{(p/m)\beta_{n,p,m}} = 1 + o_p(1). \quad (\text{A.29})$$

So we have

$$\begin{aligned} \gamma_n \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{-1/2} &= \frac{pm}{\sqrt{8m + 2p\beta_{n,p,m}^2}} (A_{n,p,m} + V_{n,p,m}) (1 + \beta_{n,p,m}) \\ &= \frac{pmA_{n,p,m}}{\sqrt{8m + 2p\beta_{n,p,m}^2}} \frac{A_{n,p,m} + V_{n,p,m}}{A_{n,p,m}} (1 + \beta_{n,p,m}) \\ &= \frac{pmA_{n,p,m}/\beta_{n,p,m}}{\sqrt{8m + 2p\beta_{n,p,m}^2}} (1 + \beta_{n,p,m})^2 (1 + o_p(1)) \quad (\text{by (A.28)}) \\ &= \frac{p}{\sqrt{8m + 2p\beta_{n,p,m}^2}} (1 + \beta_{n,p,m})^2 (1 + o_p(1)) \quad (\text{by (A.29)}) \\ &= \alpha_{n,p,m} (1 + o_p(1)), \end{aligned} \quad (\text{A.30})$$

360 where $\alpha_{n,p,m} = p (8m + 2p\beta_{n,p,m}^2)^{-1/2} (1 + \beta_{n,p,m})^2$.

Also,

$$\begin{aligned}
0 &< \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left(\frac{8}{mp^2} + \frac{2}{m^2 p} \beta_{n,p,m}^2 \right)^{-1/2} \\
&\leq \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left(\frac{2}{m^2 p} \beta_{n,p,m}^2 \right)^{-1/2} \\
&= \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left\{ \frac{2}{p^3} \left(\frac{p}{m} \beta_{n,p,m} \right)^2 \right\}^{-1/2} \\
&= \sqrt{\frac{2}{p}} \left(\ln \frac{n}{m} \right)^2 \left(2 \ln \frac{n}{m} (1 + o(1)) \right)^{-1} \\
&= \frac{1}{\sqrt{2p}} \ln \frac{n}{m} (1 + o(1)) = o(1),
\end{aligned}$$

362 which together with (A.30) implies that

$$\begin{aligned}
&\frac{\alpha_{n,p,m}}{\gamma_n} \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \\
&= \left\{ \frac{\alpha_{n,p,m}}{\gamma_n} \left(\frac{8}{mp^2} + \frac{2}{m^2 p} \beta_{n,p,m}^2 \right)^{1/2} \right\} \left\{ \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left(\frac{8}{mp^2} + \frac{2}{m^2 p} \beta_{n,p,m}^2 \right)^{-1/2} \right\} \\
&= (1 + o_p(1)) o(1) = o_p(1). \tag{A.31}
\end{aligned}$$

It follows from (A.27), (A.30), and (A.31) that

$$\begin{aligned}
&\alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \\
&\stackrel{d}{=} \frac{\alpha_{n,p,m}}{\gamma_n} \left(\frac{8}{mp^2} + \frac{2}{m^2 p} \beta_{n,p,m}^2 \right)^{1/2} \{c_{1,n} Z_{1,n} + c_{2,n} Z_{2,n} (1 + o_p(1))\} - \frac{\alpha_{n,p,m}}{\gamma_n} \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \\
&= (1 + o_p(1)) \{c_{1,n} Z_{1,n} + c_{2,n} Z_{2,n} (1 + o_p(1))\} + o_p(1). \tag{A.32}
\end{aligned}$$

364 Note that $c_{1,n}$ and $c_{2,n}$ are constants (depending on n, p_n, m_n), which satisfy

$c_{1,n}^2 + c_{2,n}^2 = 1$. By Lemma A7(iii),

$$-Z_{1,n} = \sqrt{\frac{m}{8}} \left\{ -p A_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 - \frac{2}{p} \left(\ln \frac{n}{m} \right)^2 \right\} \xrightarrow{d} N(0, 1).$$

366 By (A.22), $Z_{2,n} \xrightarrow{d} N(0, 1)$. Note that $Z_{1,n}$ and $Z_{2,n}$ are independent (since $A_{n,p,m}$ and $V'_{2,n}$ are independent). We have

$$c_{1,n}Z_{1,n} + c_{2,n}Z_{2,n} \xrightarrow{d} N(0, 1),$$

368 which together with (A.32) implies that

$$\alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \xrightarrow{d} N(0, 1).$$

The proof is complete.

370 **Proof of Theorem 4.** Part (i) follows immediately from Theorem 2. To prove part (ii), we have by part (i) and Theorem 3 that

$$\begin{aligned} & 2\alpha_{n,p,m} \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \left(\rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \right) \\ &= \frac{2\sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}}}{\rho_{n,p,m} + \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}}} \alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \xrightarrow{d} N(0, 1), \end{aligned}$$

372 completing the proof. □

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